

## Oscillator strength

Oscillator strength is a very important parameter in spectroscopy. It is defined in a dimensionless form and gives the probability of absorption or emission of radiation in transitions between energy levels of an atom or a molecule.

$$f_{fi} = \frac{2m\omega_{fi}}{3\hbar} |\langle f | \vec{r} | i \rangle|^2 \quad \omega_{fi} = \frac{E_f - E_i}{\hbar}$$

Where the dipole matrix elements are

$$|\langle f | \vec{r} | i \rangle|^2 \equiv |\langle f | x | i \rangle|^2 + |\langle f | y | i \rangle|^2 + |\langle f | z | i \rangle|^2$$

Oscillator strengths obey the so-called Thomas-Reiche-Kuhn sum rule

$$\sum_f f_{fi} = 1$$

Obviously  $f_{if} < 1$

The above sum rule can be derived as follows

$$\langle \dot{i} | [x, p_x] | \dot{i} \rangle = i\hbar \quad \leftarrow \text{basic commutator relation}$$

We may write it as

$$\sum_f \langle \dot{i} | x | f \rangle \langle f | p_x | i \rangle - \langle \dot{i} | p_x | f \rangle \langle f | x | i \rangle = i\hbar$$

Now, if the Hamiltonian is such that

$$H = \frac{\vec{p}^2}{2m} + V(\vec{r}) = \frac{p_x^2 + p_y^2 + p_z^2}{2m} + V(\vec{r}) \quad \text{we can}$$

represent  $\vec{p}'$  as

$$\vec{p}_x = \frac{im}{\hbar} [H, \vec{x}]$$

$$\begin{aligned} \text{(check: } [x, p_x^2] &= [x, p_x] p_x + p_x [x, p_x] \\ &= 2i\hbar p_x \end{aligned}$$

$$\text{Hence, } \langle f | p_x | i \rangle = \frac{i\hbar}{\hbar} \langle f | (H_x - xH) | i \rangle$$

Because both  $|f\rangle$  and  $|i\rangle$  are eigenfunctions of  $H$   
we obtain

$$\langle f | p_x | i \rangle = -\frac{i\hbar}{\hbar} (E_i - E_f) \langle f | x | i \rangle$$

or simply

$$\langle f | p_x | i \rangle = +i\omega_{fi} \langle f | x | i \rangle$$

With that

$$\sum_f \langle i | x | f \rangle \langle f | p_x | i \rangle - \langle i | p_x | f \rangle \langle f | x | i \rangle = i\hbar$$

becomes

$$\sum_f [i\omega_{fi} |\langle f | x | i \rangle|^2 + i\omega_{fi} |\langle f | x | i \rangle|^2] = i\hbar$$

$$\text{or } \sum_f \frac{2i\omega_{fi}}{\hbar} |\langle f | x | i \rangle|^2 = 1$$

The same result follows with  $x$  replaced by  
 $y$  or  $z$ . When we combine them we get

$$\sum_f \frac{2i\omega}{3\hbar} |\langle f | \vec{F} | i \rangle|^2 = \sum_f f_{fi} = 1$$

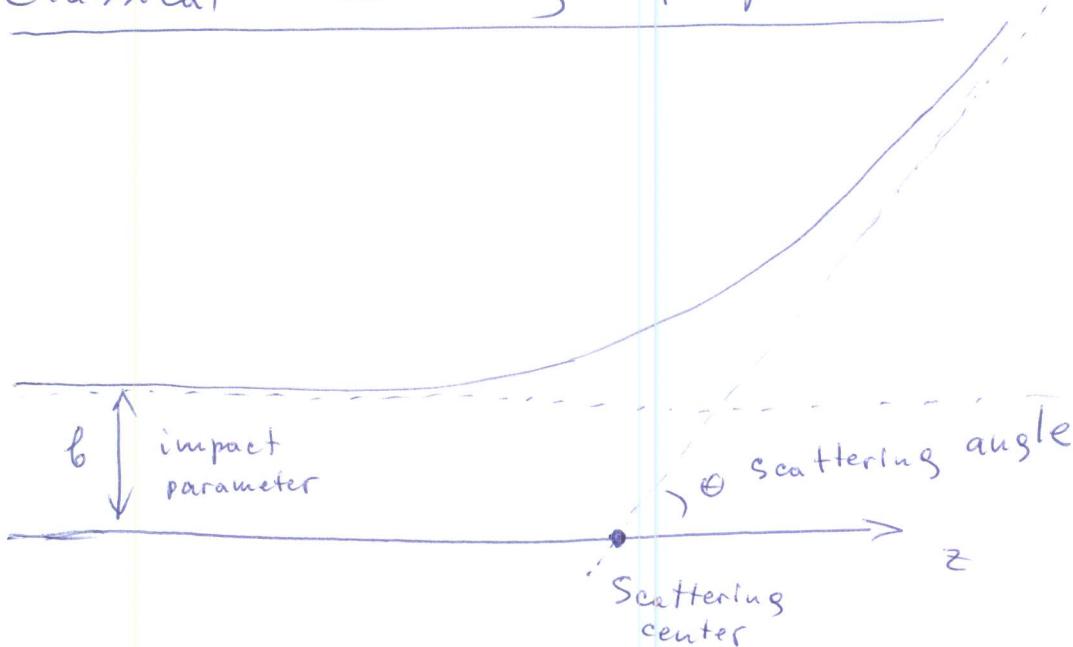
For an atom with  $Z$  electrons we make a replacement

$$\langle f | \vec{F} | i \rangle \Rightarrow \langle f | \sum_{i=1}^Z \vec{r}_i | i \rangle$$

and then

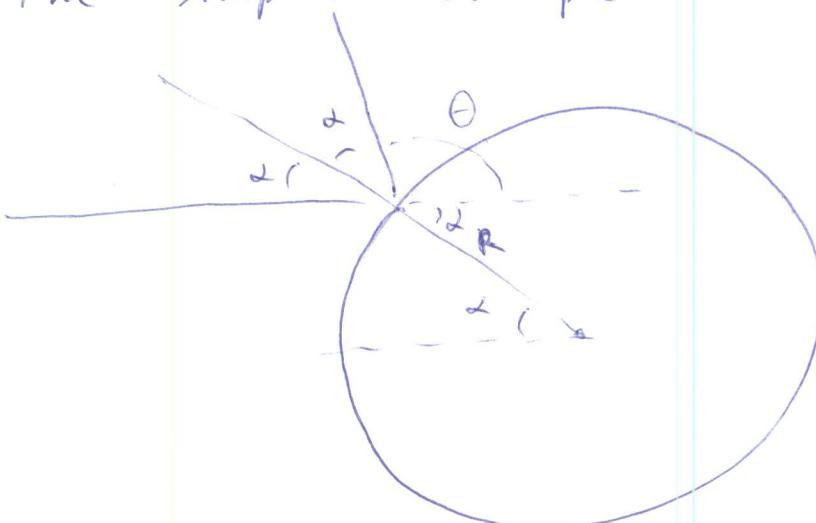
$$\sum_f f_{fi} = Z$$

# Classical scattering of particles



Let us consider a particle incident on some scattering center. The particle has some initial energy,  $E$ , and impact parameter,  $b$ . Eventually it scatters at some scattering angle  $\theta$ . For simplicity we will assume that the target's potential is spherically symmetric. Then the trajectory lies in a single plane. Our task is to compute  $\theta$  given  $b$ .

The simplest example is a hard-sphere potential



$$b = R \sin \theta$$

$$\theta = \pi - 2\alpha$$

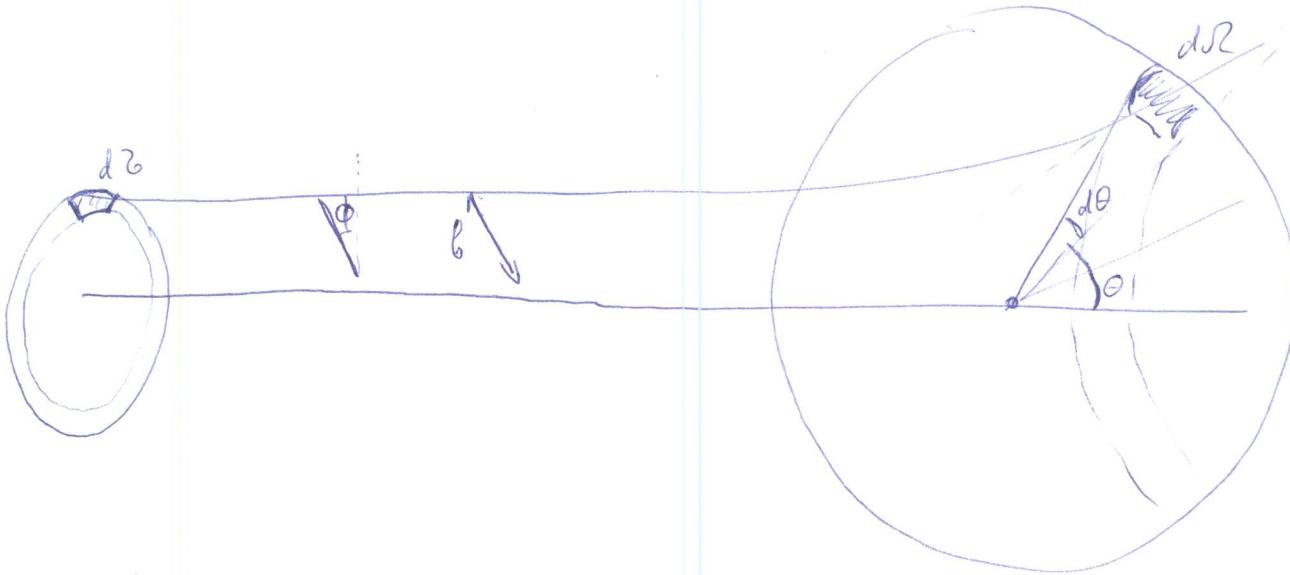
$$b = R \sin\left(\frac{\pi}{2} - \frac{\theta}{2}\right) = R \cos\theta$$

$$\theta = \begin{cases} 2 \arccos \frac{b}{R}, & b < R \\ 0, & b \geq R \end{cases}$$

## Differential cross section

$$D(\theta) = \frac{d\sigma}{d\Omega}$$

$d\sigma$  - cross-sectional area  
 $d\Omega$  - solid angle



$$d\sigma = D(\theta) d\Omega$$

$$d\sigma = b db d\phi$$

$$d\Omega = \sin\theta d\theta d\phi$$

$$D(\theta) = \frac{b}{\sin\theta} \left| \frac{db}{d\theta} \right|$$

For a hard sphere

$$\frac{db}{d\theta} = -\frac{1}{2} R \sin \frac{\theta}{2}$$

$$\text{and } D(\theta) = \frac{R \cos \frac{\theta}{2}}{\sin \theta} \left( \frac{R \sin \frac{\theta}{2}}{2} \right) = \frac{R^2}{4}$$

The total cross-section is the integral of  $D(\theta)$  over all solid angles:

$$\sigma = \int D(\theta) d\Omega$$

For a hard sphere

$$\sigma = \frac{R^2}{4} \int d\Omega = \pi R^2$$

If we have a beam of incident particles with uniform intensity (luminosity)

$L$  - # particles per unit area per unit time

then the number of particles going through area  $d\Omega$  (and scattered into solid angle  $d\Omega$ ) per unit time is

$$dN = L d\Omega = L D(\theta) d\Omega$$

So

$$D(\theta) = \frac{1}{L} \frac{dN}{d\Omega}$$

This is taken as the definition of the differential cross section because it makes reference only to quantities easily measured in the experiment.

As a model to practice let us consider a Coulomb system of charges  $q_1$  and  $q_2$  and masses  $m_1=m$  and  $M_2=\infty$ .

Conservation of energy

$$\text{where } V(r) = \frac{q_1 q_2}{4\pi \epsilon_0 r}$$

Conservation of angular

$$E = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) + V(r)$$

$$\text{momentum: } L = m r^2 \dot{\phi} \quad \text{so } \dot{\phi} = \frac{L}{m r^2}$$

$$\dot{r}^2 + \frac{L^2}{m r^2} = \frac{2}{m} (E - V)$$

$$\begin{aligned} \dot{r} &= \frac{dr}{dt} = \frac{dr}{du} \frac{du}{d\phi} \frac{d\phi}{dt} = \\ &= -\frac{1}{m} \frac{du}{d\phi} \frac{L}{m} u^2 = -\frac{L}{m} \frac{du}{d\phi} \end{aligned}$$

$$\text{let } u = \frac{1}{r} \quad \text{then}$$

