

Partial wave analysis

For a spherically symmetric potential $V(r)$ we can use the separation of variables $\psi(r, \theta, \phi) = R(r) Y_l^m(\theta, \phi)$ and the radial equation for $R(r) = \frac{u(r)}{r}$ looks as follows

$$-\frac{\hbar^2}{2m} \frac{d^2 u}{dr^2} + \left[V(r) + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2} \right] u = E u$$

At $r \rightarrow \infty$ $V(r) \rightarrow 0$ and we get

$$\frac{d^2 u}{dr^2} = -k^2 u$$

with the general solution $u(r) = C e^{ikr} + D e^{-ikr}$. For the scattered wave the incoming term $D e^{-ikr}$ must vanish. Thus, $D = 0$. Then

$$\text{at } r \rightarrow \infty \quad R(r) \rightarrow \frac{e^{ikr}}{r}$$

If we assume that $V(r)$ is a short-range potential (i.e. $V(r)$ decays faster than $\frac{1}{r^2}$ at $r \rightarrow \infty$) we can build up the next approximation, where we neglect $V(r)$ at $r \rightarrow \infty$ but keep $\frac{l(l+1)}{r^2}$ term:

$$\frac{d^2 u}{dr^2} - \frac{l(l+1)}{r^2} u = -k^2 u$$

or

$$r^2 \frac{d^2 R}{dr^2} + 2r \frac{dR}{dr} + [k^2 r^2 - l(l+1)] R = 0$$

The solutions of this equation are called spherical Bessel and spherical Neuman functions. The above equation can be reduced to the standard Bessel equation by a substitution $R(r) = \frac{Z(kr)}{\sqrt{kr}} = \frac{Z(x)}{\sqrt{x}}$

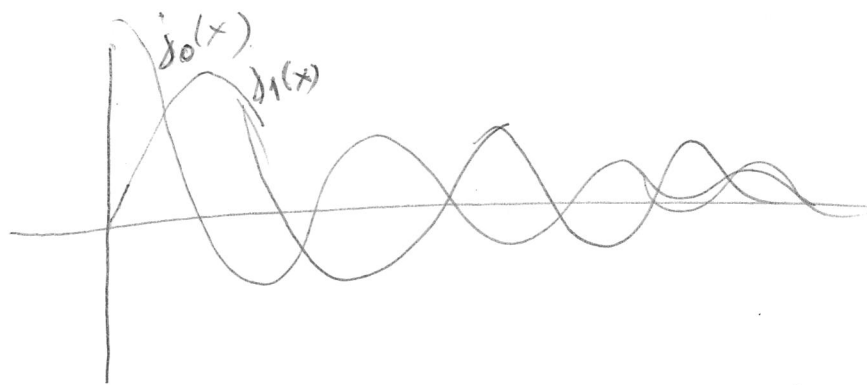
$$x^2 z'' + x z' + [x^2 - (\ell + \frac{1}{2})^2] z = 0$$

$$R(r) = A \frac{J_{\ell + \frac{1}{2}}(kr)}{\sqrt{kr}} + B \frac{N_{\ell + \frac{1}{2}}(kr)}{\sqrt{kr}} = \underbrace{A' j_\ell(kr) + B' h_\ell(kr)}_{\text{general solution}}$$

$$j_0(x) = \frac{\sin x}{x} \quad h_0(x) = -\frac{\cos x}{x}$$

$$j_1(x) = \frac{\sin x}{x^2} - \frac{\cos x}{x} \quad h_1 = -\frac{\cos x}{x^2} - \frac{\sin x}{x}$$

$$j_2(x) = \left(\frac{3}{x^3} - \frac{1}{x}\right) \sin x - \frac{3}{x^2} \cos x$$



Neither j_ℓ nor h_ℓ represent an outgoing or incoming wave. Similarly to the familiar transformation

$$A \cos x + B \sin x \rightarrow A' e^{ix} + B' e^{-ix}$$

we can introduce new linear combinations called spherical Hankel functions of the first and second kind

$$h_\ell^{(2)}(x) \equiv j_\ell(x) + i h_\ell(x)$$

$$h_0^{(1)} = -i \frac{e^{ix}}{x}$$

$$h_0^{(2)} = i \frac{e^{-ix}}{x}$$

$$h_1^{(1)} = \left(-\frac{i}{x^2} - \frac{1}{x}\right) e^{ix}$$

$$h_1^{(2)} = \left(\frac{i}{x^2} - \frac{1}{x}\right) e^{-ix}$$

At large r $h_\ell^{(1)}(kr)$ goes like $\frac{e^{ikr}}{r}$ whereas

$$h_\ell^{(2)}(kr) \rightarrow \frac{e^{-ikr}}{r}$$

With that the wave function outside the scattering region ($V(r) = 0$) can be expanded as

$$\psi(r, \theta, \phi) = A \left\{ \underbrace{e^{ikz}}_{\text{incident wave}} + \underbrace{\sum_{\ell m} c_{\ell m} h_{\ell}^{(1)}(kr) Y_{\ell}^m(\theta, \phi)}_{\text{scattered wave}} \right\}$$

In our case there is no dependence on ϕ due to the spherical symmetry of $V(r)$ so only terms with $m=0$ survive (recall $Y_{\ell}^m \sim e^{im\phi}$). Given that $Y_{\ell}^0(\theta, \phi) = \sqrt{\frac{2\ell+1}{4\pi}} P_{\ell}(\cos\theta)$ where P_{ℓ} are Legendre polynomials

we obtain

$$\psi(r, \theta) = A \left\{ e^{ikz} + \sum_{\ell=0}^{\infty} i^{\ell+1} (2\ell+1) a_{\ell} h_{\ell}^{(1)}(kr) P_{\ell}(\cos\theta) \right\}$$

where we defined the coefficients a_{ℓ} in such a way that $c_{\ell 0} = i^{\ell+1} \kappa \sqrt{4\pi(2\ell+1)} a_{\ell}$

For very large r $h_{\ell}^{(1)}(kr) \rightarrow (-i)^{\ell+1} \frac{e^{ikr}}{kr}$, so

$$\psi(r, \theta) \rightarrow A \left\{ e^{ikz} + f(\theta) \frac{e^{ikr}}{r} \right\}$$

with

$$f(\theta) = \sum_{\ell=0}^{\infty} (2\ell+1) a_{\ell} P_{\ell}(\cos\theta)$$

The differential cross section is then

$$\frac{d\sigma}{d\Omega} = |f(\theta)|^2 = \sum_{\ell} \sum_{\ell'} (2\ell+1)(2\ell'+1) a_{\ell}^* a_{\ell'} P_{\ell}(\cos\theta) P_{\ell'}(\cos\theta)$$

while the total cross section becomes (if we use the orthogonality of P_{ℓ} : $\int_0^{\pi} P_{\ell}(\cos\theta) P_{\ell'}(\cos\theta) \sin\theta d\theta = \frac{2}{2\ell+1} \delta_{\ell\ell'}$)

$$G = 4\pi \sum_{\ell=0}^{\infty} (2\ell+1) |a_{\ell}|^2$$

The question that remains now is how to determine coefficients a_{ℓ} called partial wave amplitudes. This is done by solving the SE in the interior region (where $V \neq 0$) and matching it to the exterior solution $\psi = A \left\{ e^{ikz} + k \sum_{\ell=0}^{\infty} i^{\ell} (2\ell+1) a_{\ell} h_{\ell}^{(1)}(kr) P_{\ell}(\cos\theta) \right\}$ using the appropriate boundary conditions. To make things easier we expand $e^{ikr} = e^{ikr \cos\theta}$ in terms of spherical Bessel functions (any nonsingular function $g(x)$ can be expanded in terms of $j_{\ell}(x)$ as the latter form a complete set). It is known that

$$e^{ikz} = \sum_{\ell=0}^{\infty} i^{\ell} (2\ell+1) j_{\ell}(kr) P_{\ell}(\cos\theta) \quad \text{--- Rayleigh formula}$$

with that we can write the exterior solution as

$$\psi(r, \theta) = A \sum_{\ell=0}^{\infty} i^{\ell} (2\ell+1) \left[j_{\ell}(kr) + ik a_{\ell} h_{\ell}^{(1)}(kr) \right] P_{\ell}(\cos\theta)$$

To illustrate the above approach let us consider quantum scattering from a hard sphere:

$$V(r) = \begin{cases} \infty & r \leq b \\ 0 & r > b \end{cases} \quad b - \text{radius of the hard sphere}$$

Boundary condition $\psi(b, \theta) = 0$, so

$$\sum_{\ell=0}^{\infty} i^{\ell} (2\ell+1) \left[j_{\ell}(kb) + ik a_{\ell} h_{\ell}^{(1)}(kb) \right] P_{\ell}(\cos\theta) = 0 \quad \forall \theta$$

Multiplying by $P_{\ell}(\cos\theta) \sin\theta d\theta$ and integrating from 0 to π we get

$$2i^l [j_l(kb) + ik a_l h_l^{(1)}(kb)] = 0$$

and hence

$$a_l = - \frac{j_l(kb)}{ik h_l^{(1)}(kb)}$$

The total cross section is

$$\sigma = \frac{4\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \left| \frac{j_l(kb)}{h_l^{(1)}(kb)} \right|^2$$

For low energy scattering $kb \ll 1$ ($k = \frac{2\pi}{\lambda}$ - the wavelength is much greater than b). For small value of the argument

$$\frac{j_l(x)}{h_l^{(1)}(x)} = \frac{j_l(x)}{j_l(x) + i h_l(x)} \approx -i \frac{j_l(x)}{h_l(x)} \approx \frac{i}{2l+1} \left[\frac{2^l l!}{(2l)!} \right]^2 x^{2l+1}$$

(we used $j_l(x) \xrightarrow{x \rightarrow 0} \frac{2^l l!}{(2l+1)!} x^{l+1}$ $h_l(x) \xrightarrow{x \rightarrow 0} \frac{(2l)!}{2^l l!} \frac{1}{x^{l+1}}$)

With that we obtain:

$$\sigma \approx \frac{4\pi}{k^2} \sum_{l=0}^{\infty} \frac{1}{2l+1} \left[\frac{2^l l!}{(2l)!} \right]^4 (kb)^{4l+2} \approx 4\pi b^2$$

Phase shifts

The notion of phase shifts can be nicely illustrated in 1D case, when we have an impenetrable wall at $x=0$ and some localized potential near it. The incident wave $\psi_i = A e^{ix}$ ($x < -a$) and the reflected wave $\psi_r = B e^{-ix}$ ($x < -a$) [here we assume V is nonzero at $-a < x < 0$, and then $V = \infty$ at $x > 0$]. No matter what happens in the interaction region

$(-a < x < 0)$ the amplitude of the reflected wave must be the same as that of the incident wave. By conservation of probability. The only thing that can change is the phase. If $V=0$ everywhere then $B = -A$ since $\psi_{total} = \psi_i + \psi_r$ must vanish at $x=0$ (as $V(x>0) = \infty$)

$$\psi_{total} = A(e^{ikx} - e^{-ikx})$$

If the potential is not zero

$$\psi_{total} = A(e^{ikx} - e^{i(2\delta - kx)}) \quad \delta\text{'s are called phase shifts.}$$

The whole theory of elastic scattering reduces to calculating the phase shifts δ as a function of k .

In 3D case the incident wave carries no angular momentum in z -direction. Because the angular momentum is conserved (V is spherically symmetric) each partial wave scatters independently with no change in amplitude. If $V \equiv 0$ then the l -th partial wave is $\psi_{total} = A e^{ikz}$ or

$$\psi_{total}^{(e)} = A i^l (2l+1) j_l(kr) P_l(\cos\theta)$$

Meanwhile

$$j_l(x) = \frac{1}{2} [h_l^{(1)}(x) + h_l^{(2)}(x)] \approx \frac{1}{2x} [(-i)^{l+1} e^{ix} + i^{l+1} e^{-ix}] \quad x \gg 1$$

When $r \rightarrow \infty$

$$\psi_{total}^{(e)} \approx A \frac{(2l+1)}{2ikr} [e^{ikr} + (-1)^{l+1} e^{-ikr}] P_l(\cos\theta)$$

The second term in the square brackets represents

an incoming spherical wave. The first term is the outgoing wave. It picks up a phase shift δ_e

$$\psi^{(e)} \approx A \frac{(2l+1)}{2ikr} \left[e^{i(kr + 2\delta_e)} - (-1)^l e^{-ikr} \right] P_l(\cos\theta)$$

Previously we had expressed everything in terms of partial wave amplitudes a_e . Now we have expressed everything in terms of δ_e . The connection between the two is established when we consider the asymptotic form of $\psi(r, \theta) = A \left\{ e^{iuz} + \sum_{l=0}^{\infty} i^{l+1} (2l+1) a_e h_e^{(1)}(kr) P_l(\cos\theta) \right\}$

$$\psi^{(e)} \approx A \left\{ \frac{2l+1}{2ikr} \left[e^{ikr} - (-1)^l e^{-ikr} \right] + \frac{2l+1}{r} a_e e^{ikr} \right\} P_l(\cos\theta)$$

By comparing this with the above formula we find:

$$1 + 2ik a_e = e^{2i\delta_e}$$

$$\text{and } a_e = \frac{1}{2ik} \left(e^{2i\delta_e} - 1 \right) = \frac{1}{k} e^{i\delta_e} \sin \delta_e$$

It follows that

$$f(\theta) = \frac{1}{k} \sum_{l=0}^{\infty} (2l+1) e^{i\delta_e} \sin \delta_e P_l(\cos\theta)$$

and

$$\sigma = \frac{4\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_e$$